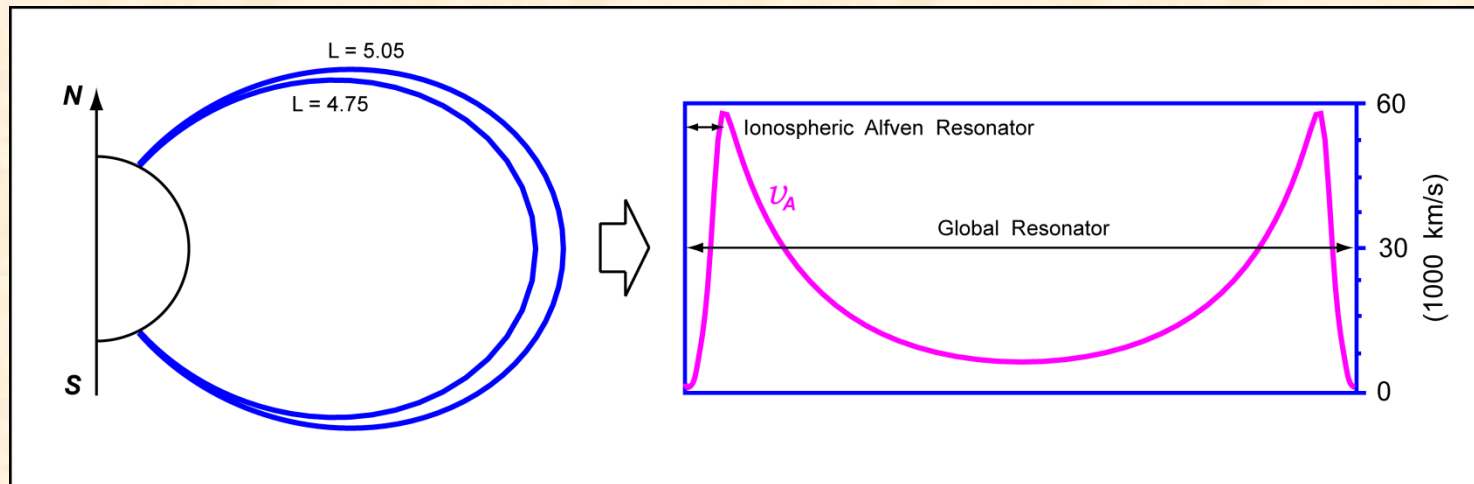


Excitation of Magnetospheric Resonators with HAARP

Anatoly V. Streltsov

Embry-Riddle Aeronautical University

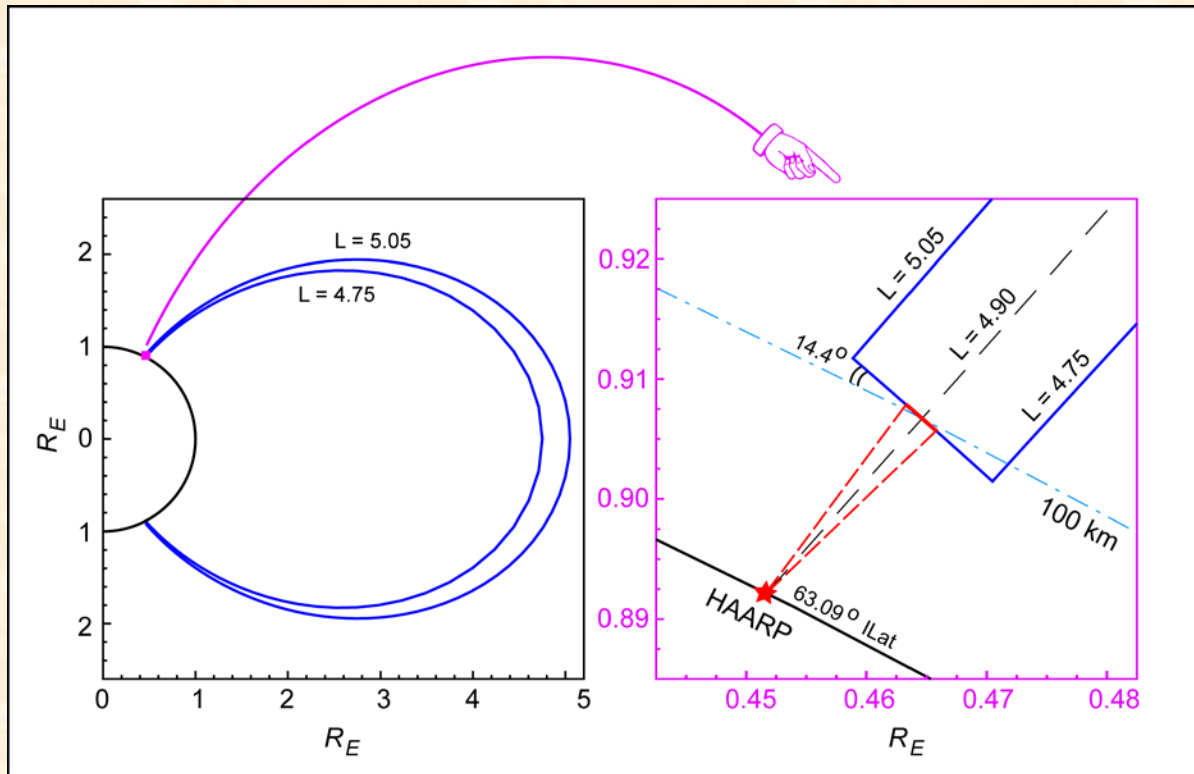
HAARP can excite ULF waves inside 1) Global Magnetospheric Resonator and 2) Ionospheric Alfvén Resonator



Mechanism of the wave generation is changing of the ionospheric conductivity by heating the ionosphere with RF waves when the electric field exists there.

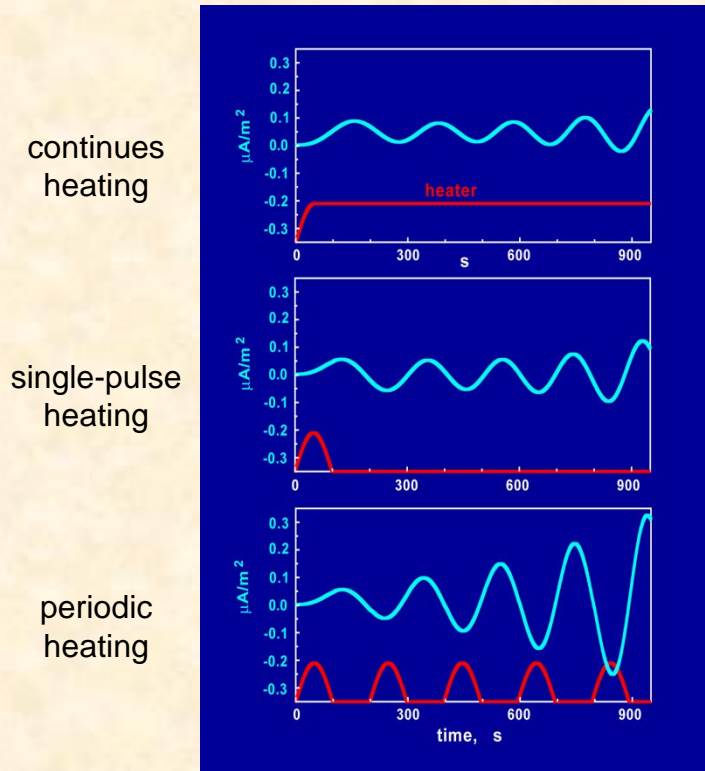
Global Resonator

HAARP Heating Experiment

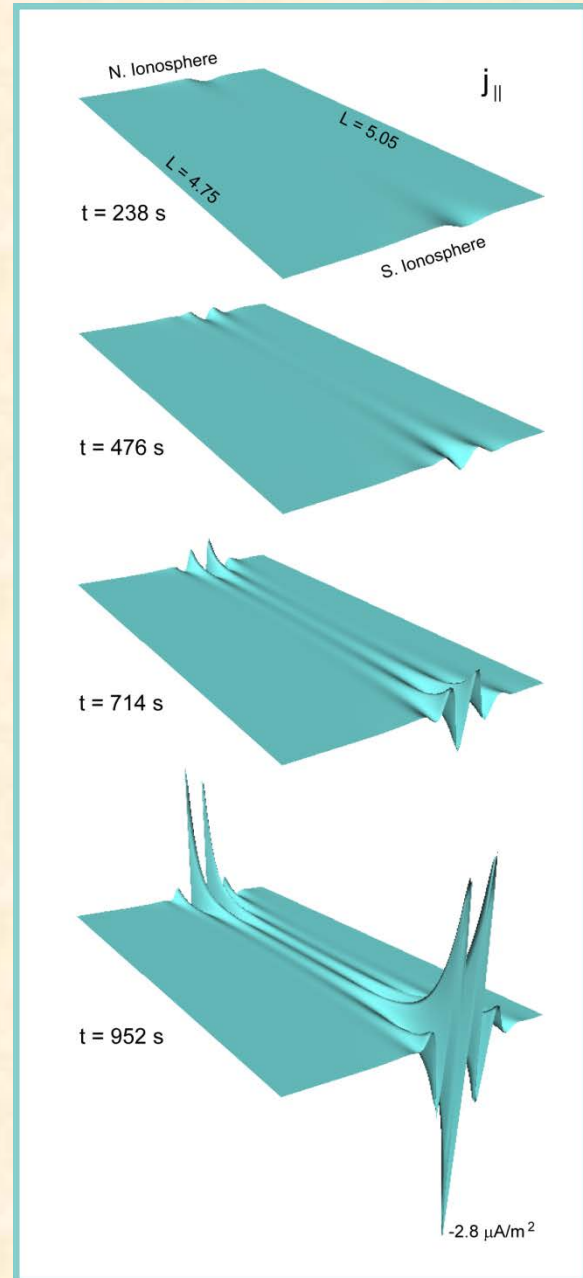


Simulations of different heating regimes for IFI excitation

$$\nabla \cdot [(\Sigma_{P0} + \mathbf{E} \Sigma_P)_{\perp 0}] = -j_{\parallel}$$



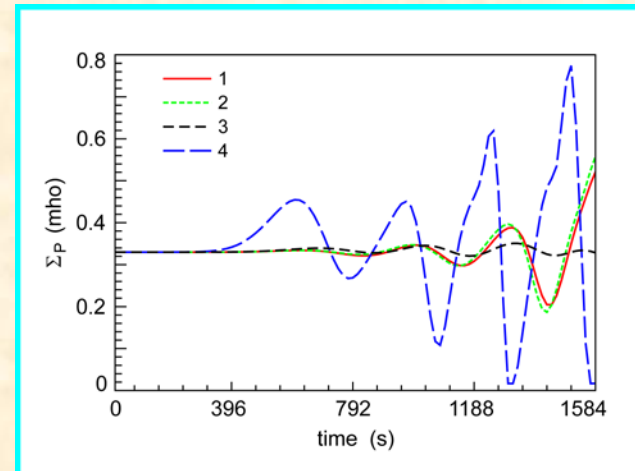
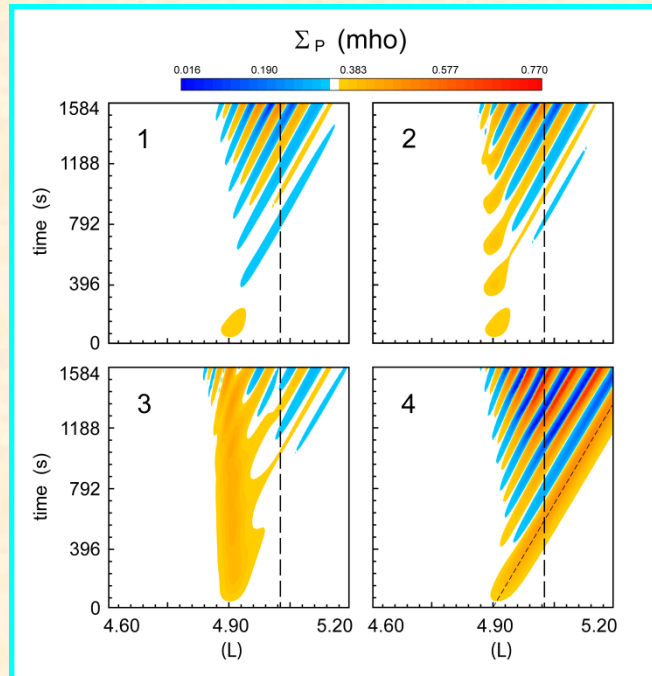
[Streltsov et al., 2005]



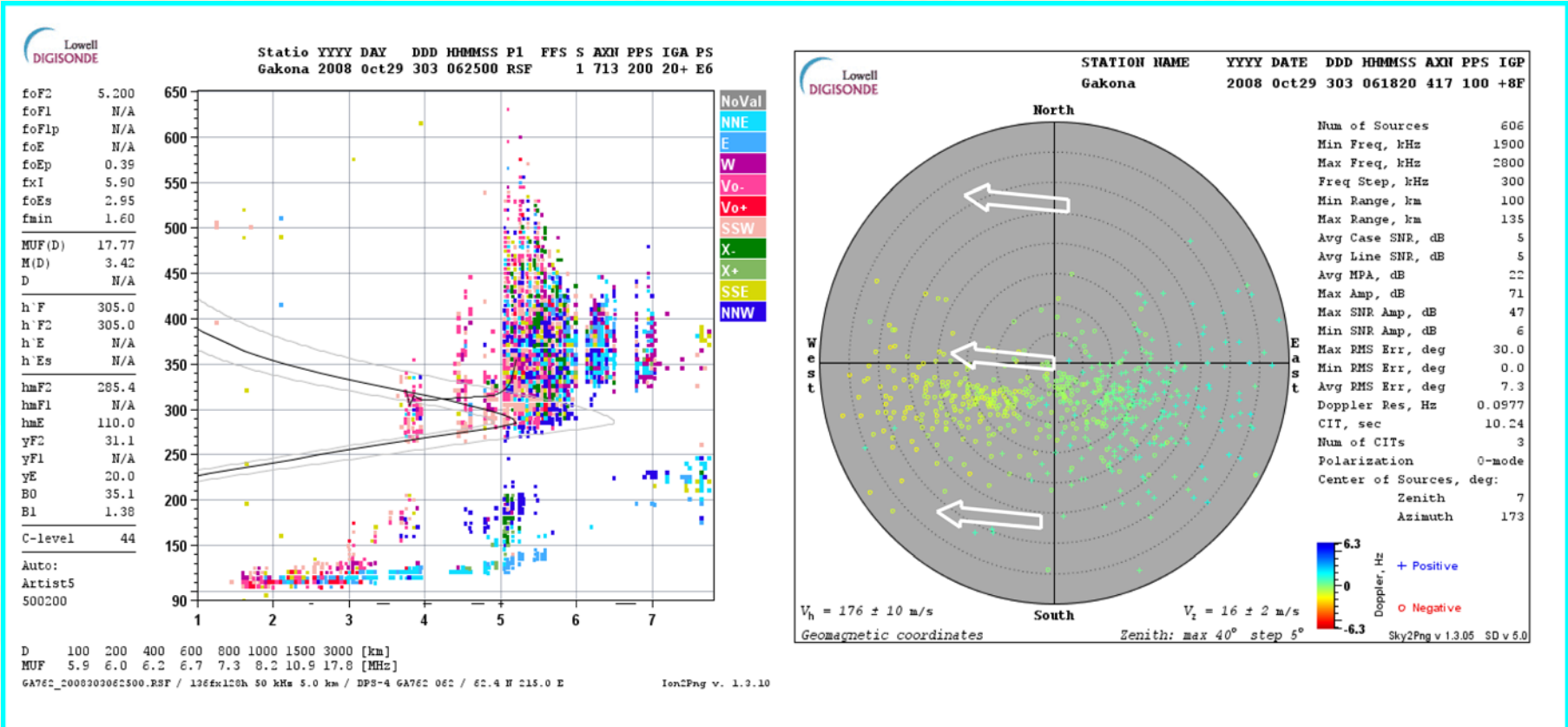
Different Heating Regimes for IFI Excitation

Old approach: heating of the same spot in the ionosphere [Streltsov *et al.*, 2005]

New approach: moving the spot with the phase velocity of the wave [Streltsov and Pedersen, 2010]

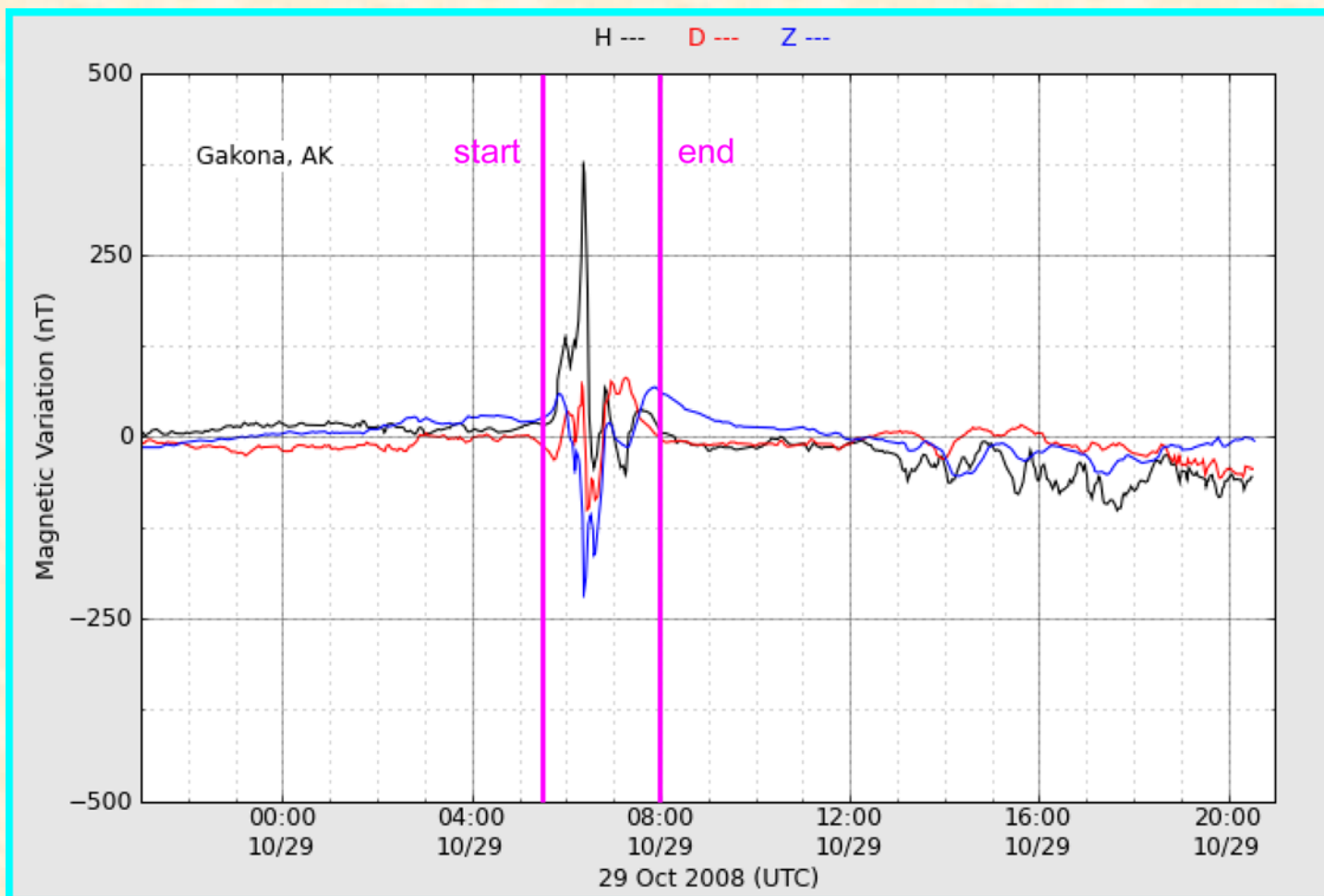


HAARP Experiment 29 Oct 2008



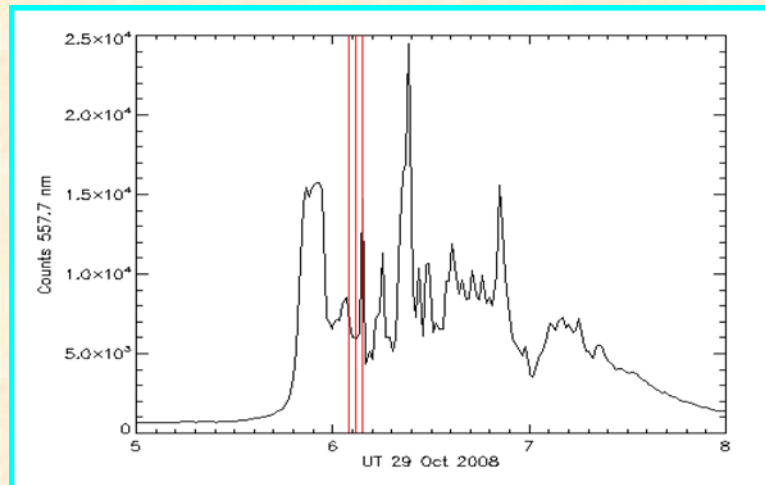
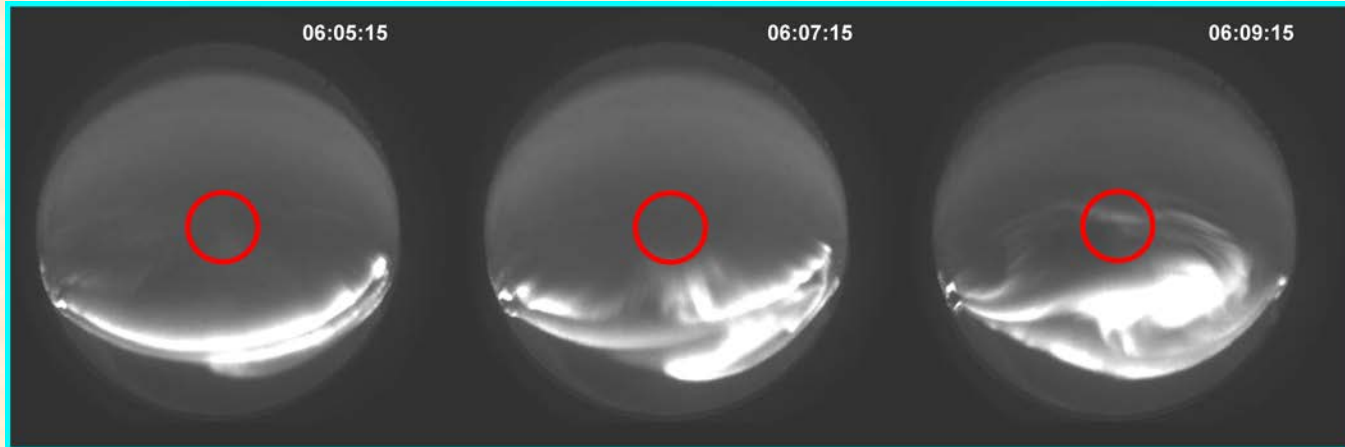
HAARP HF digisonde
 (A. Lee Snyder)

HAARP Experiment 29 Oct 2008



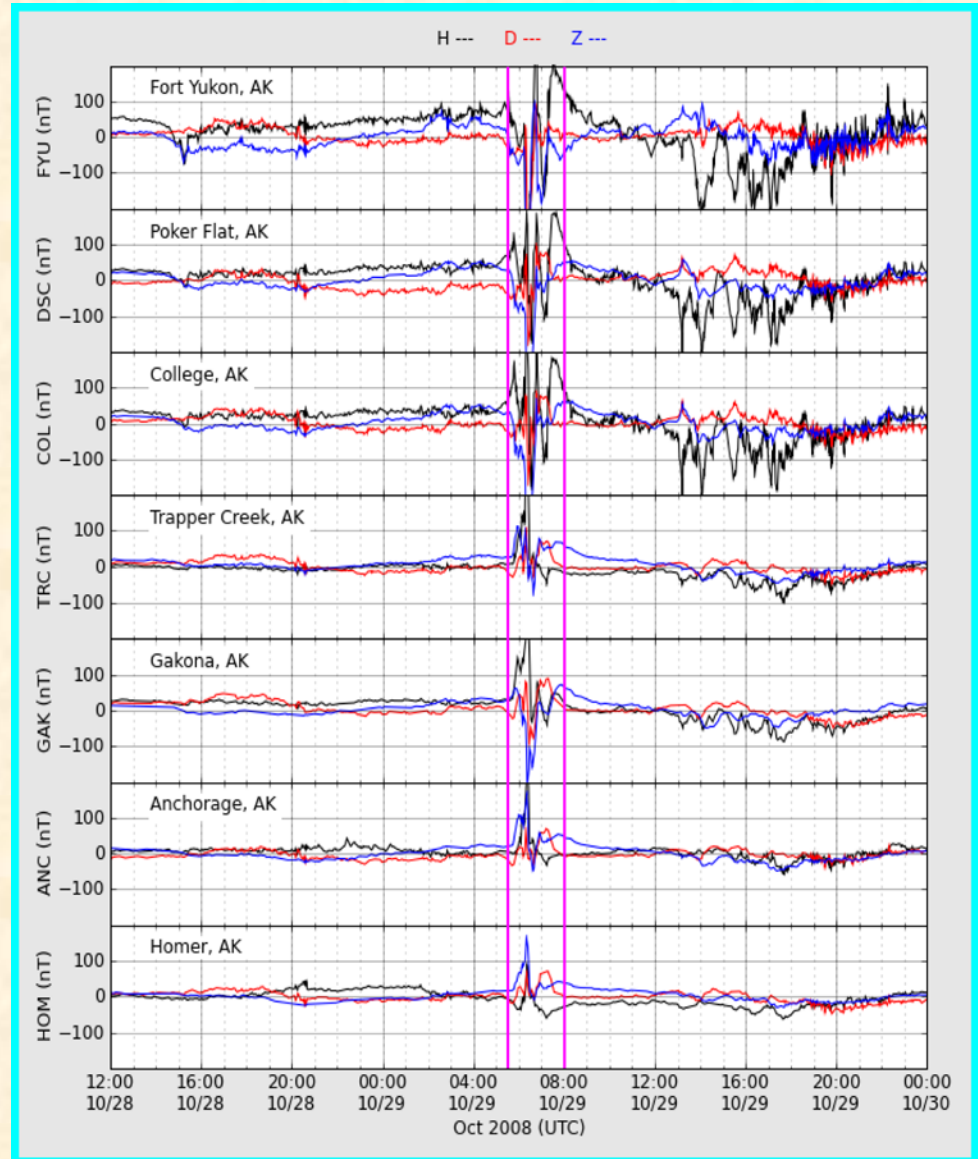
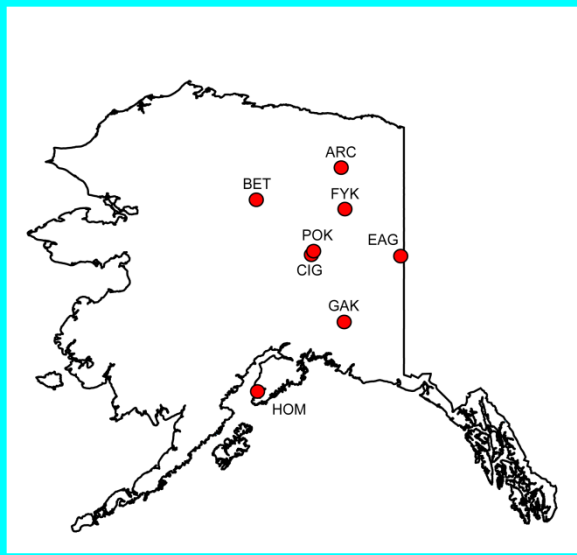
HAARP all-sky imager

(Todd Pedersen)

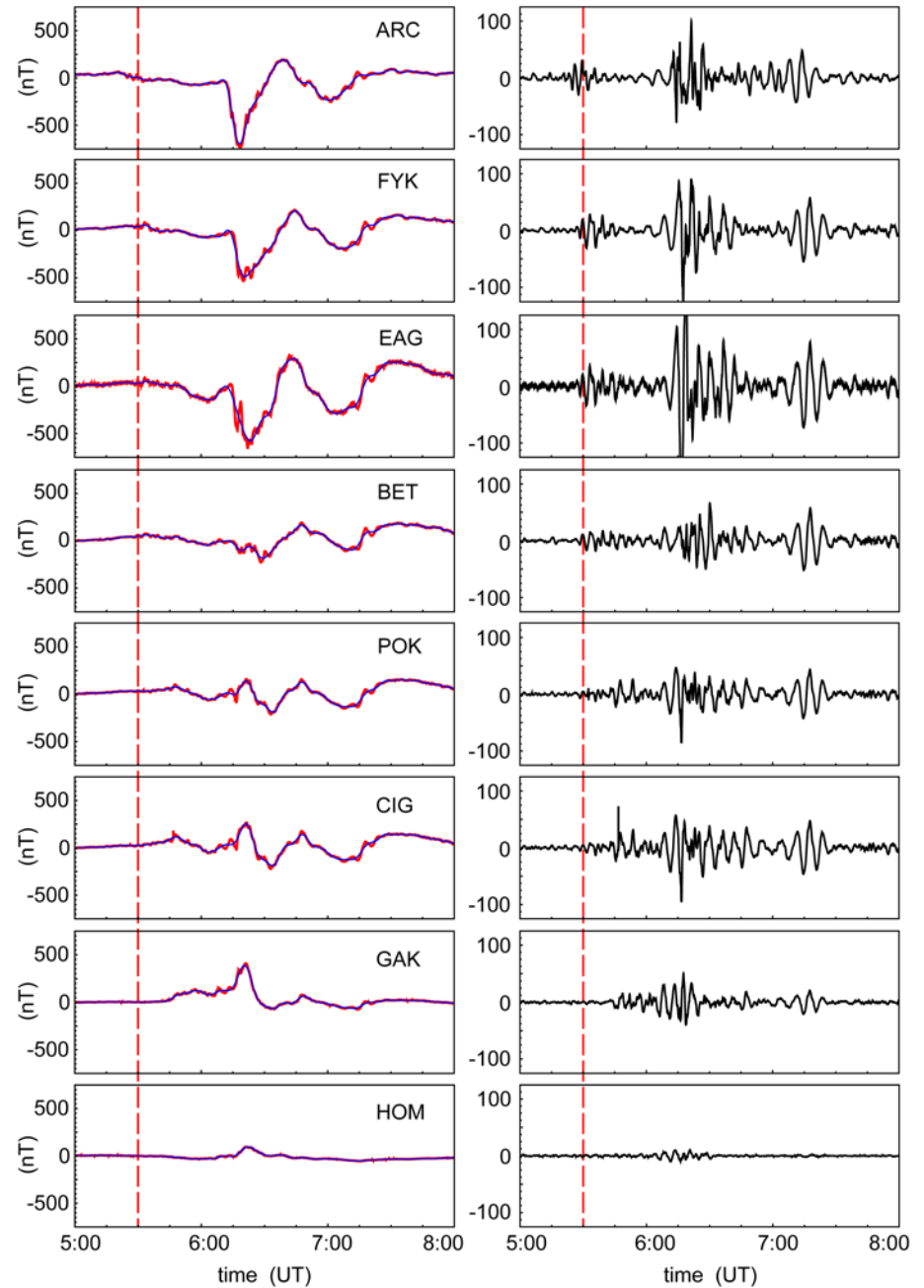
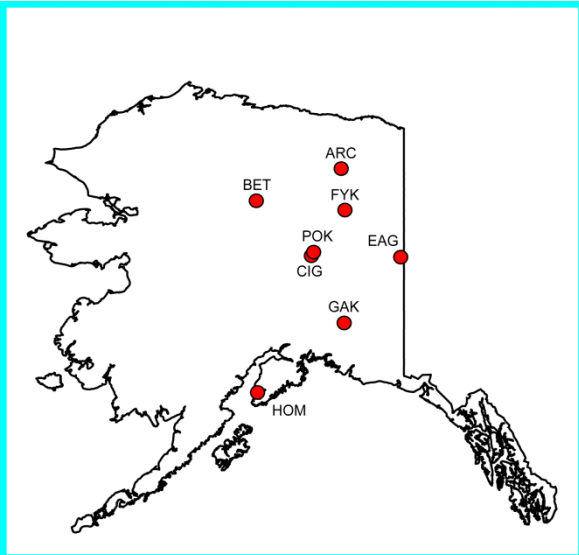


UAF/GI magnetometer array (MAGI)

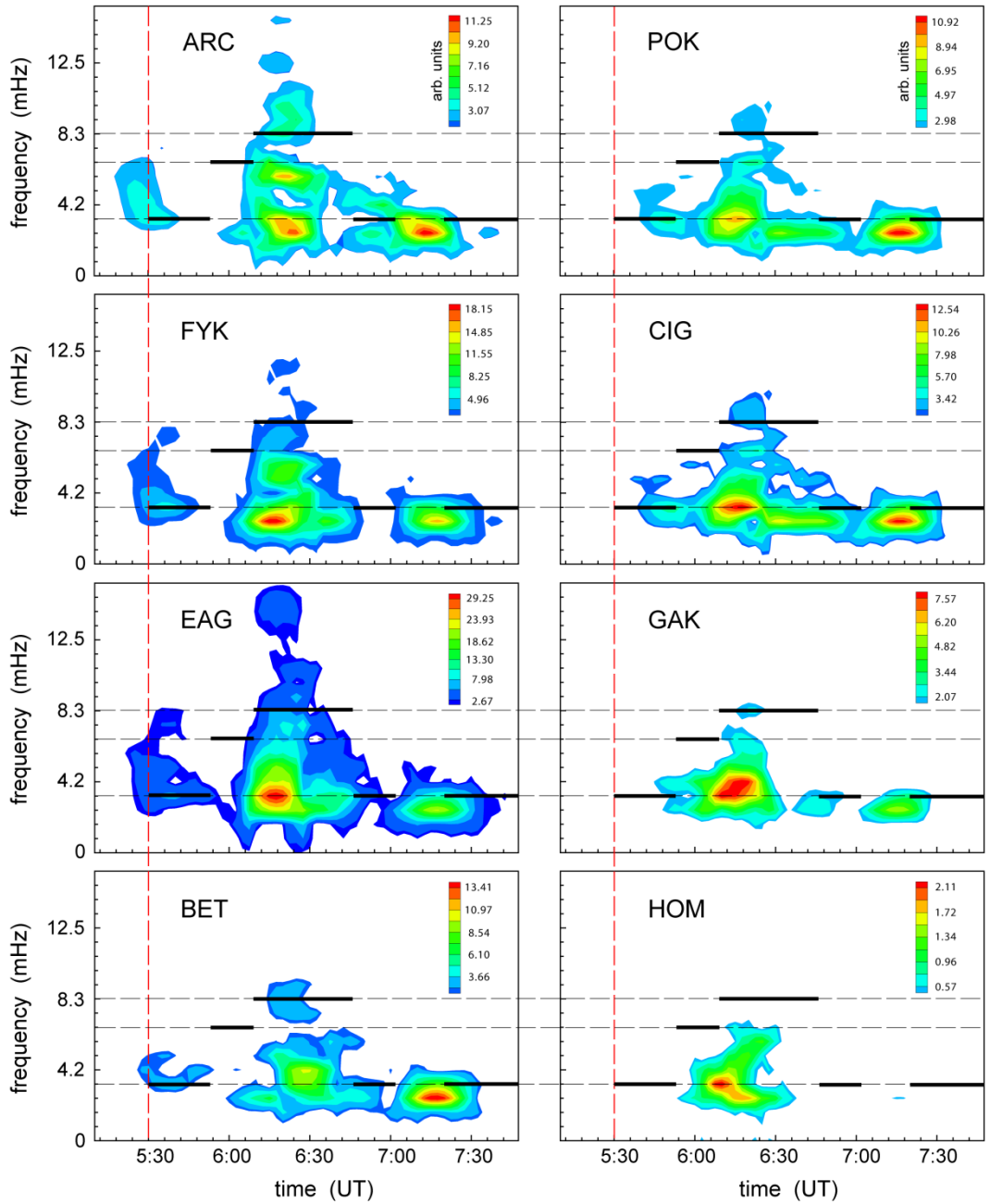
Alaska



Alaska

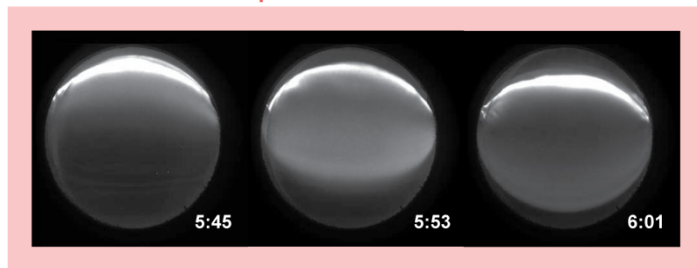
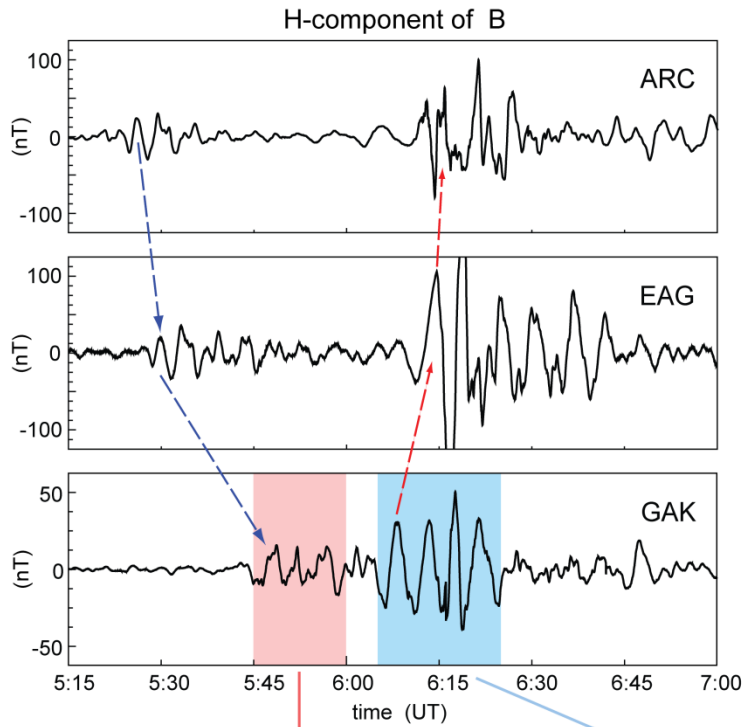


M
A
G
I



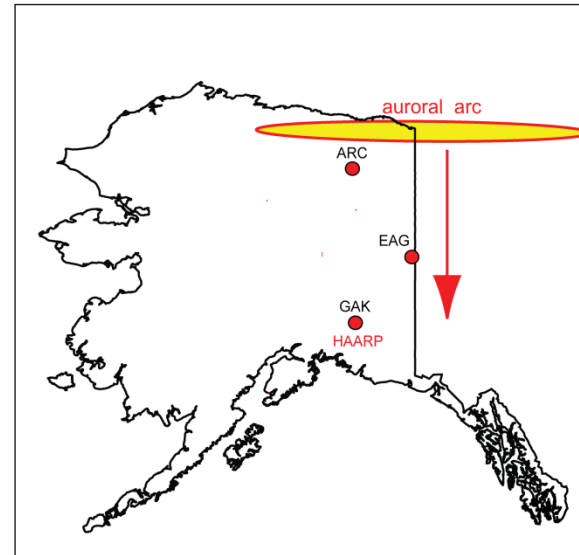
M
A
G
I

This is how it happens ...

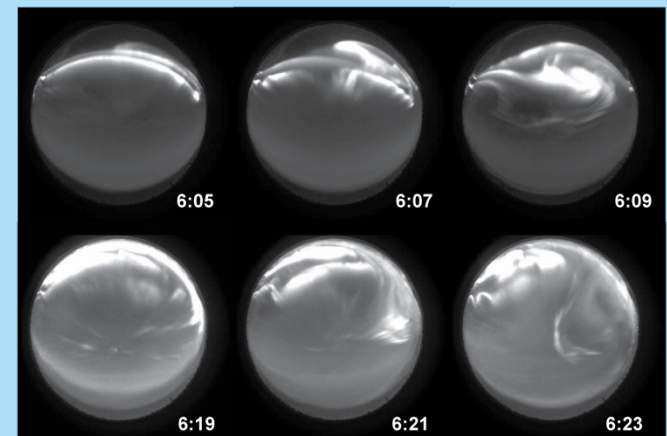


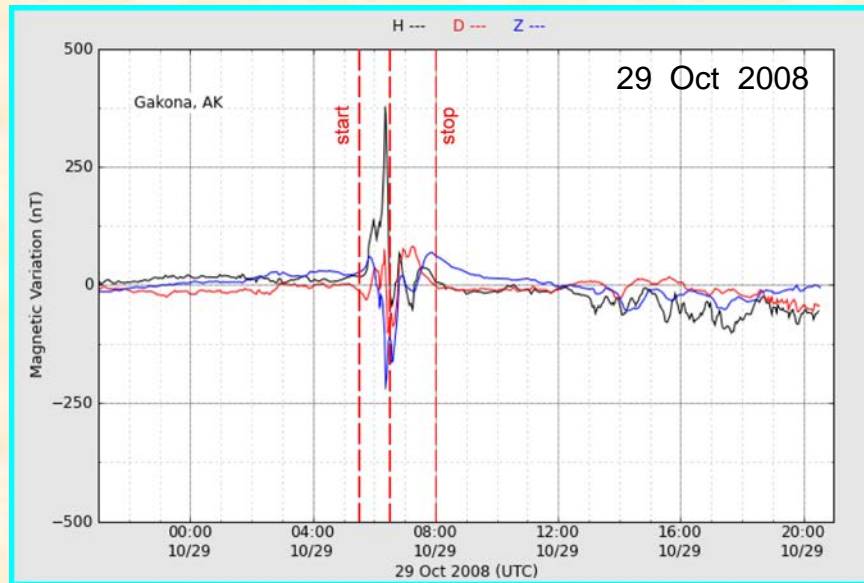
HAARP all-sky camera (557.7 nm)

Alaska (MAGI)

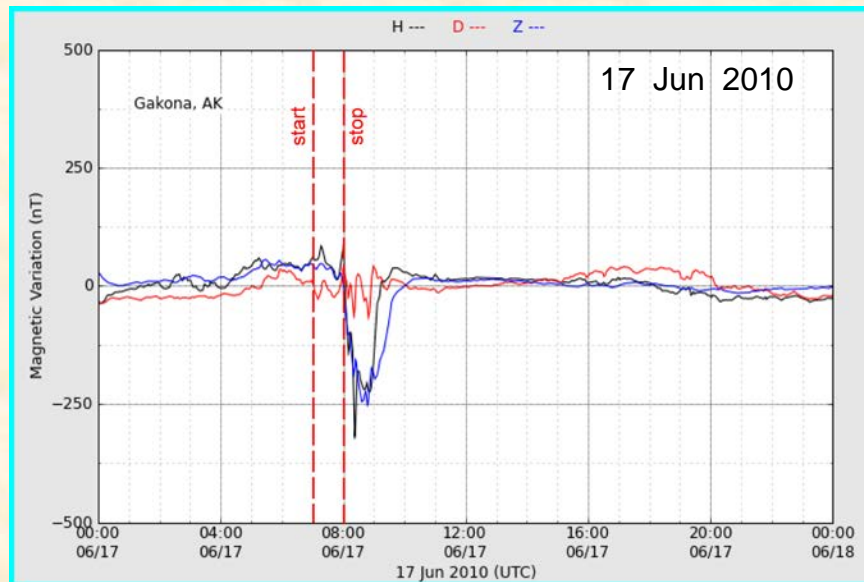


HAARP all-sky camera (557.7 nm)

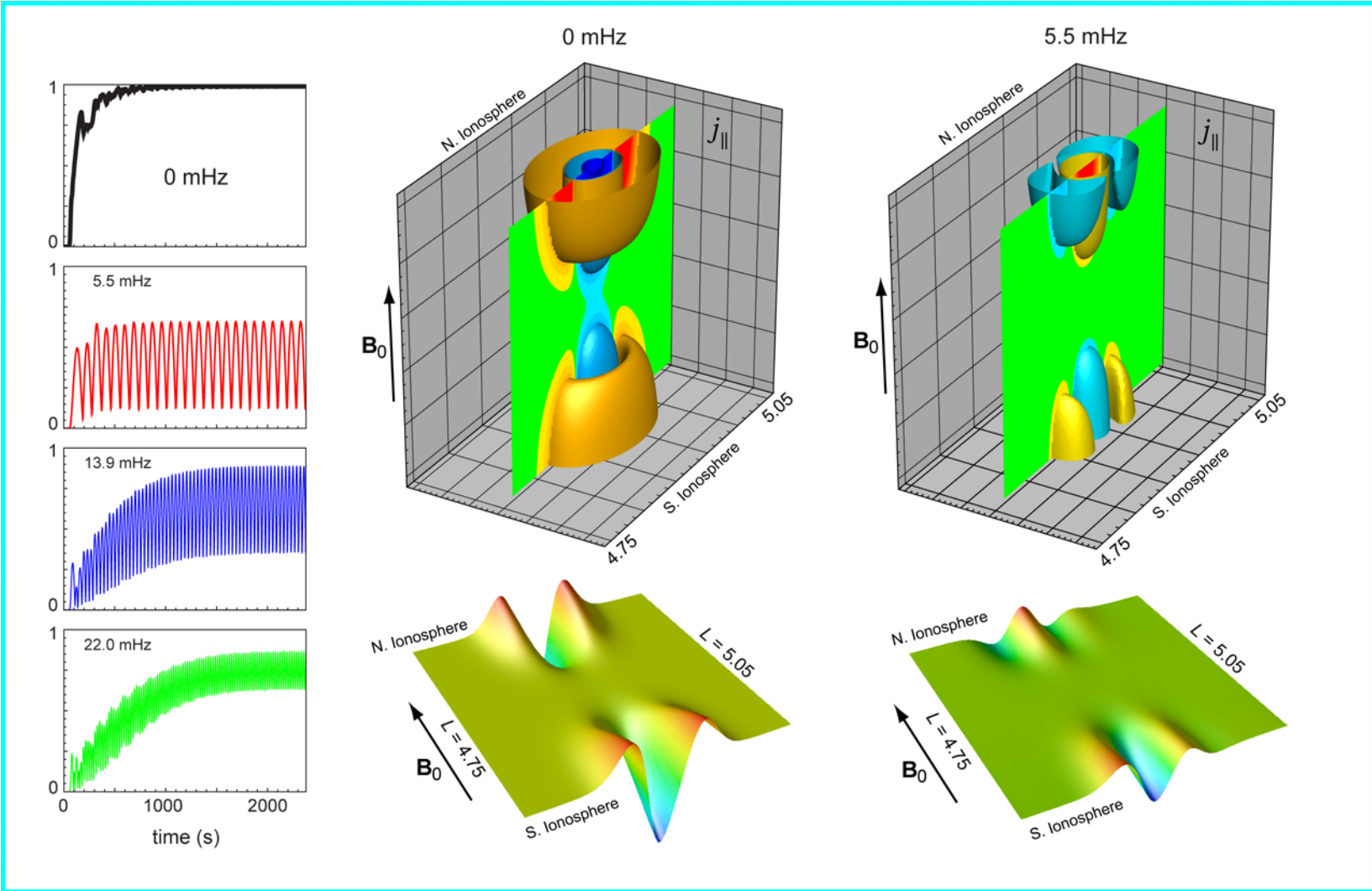




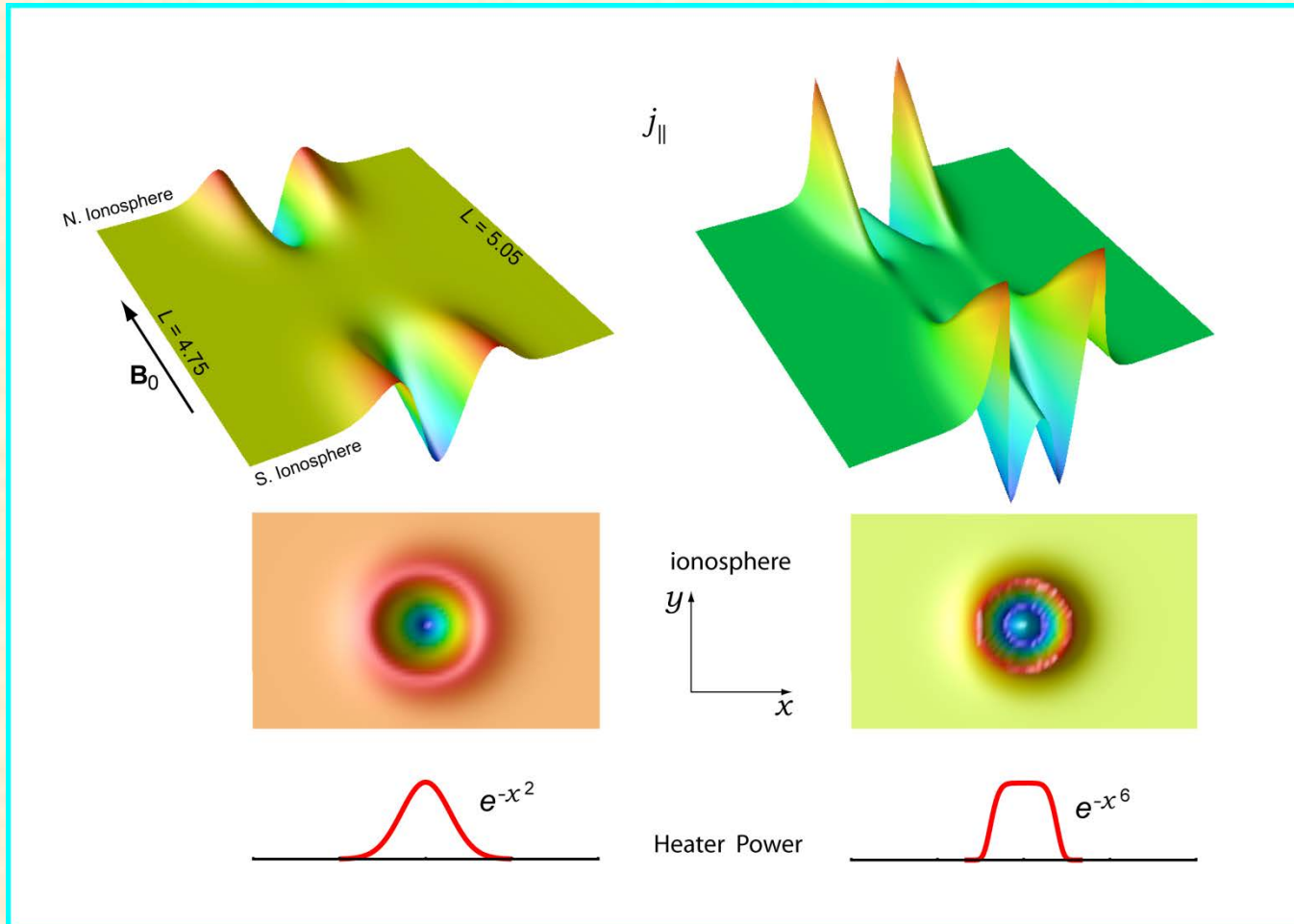
Substorms



HAARP 3D

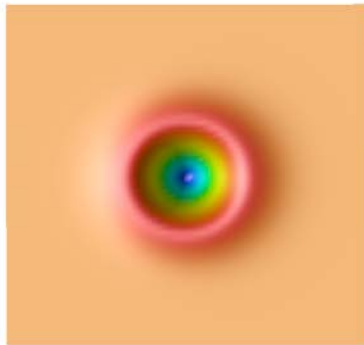


3D simulations of constant heating (high ionospheric conductivity)

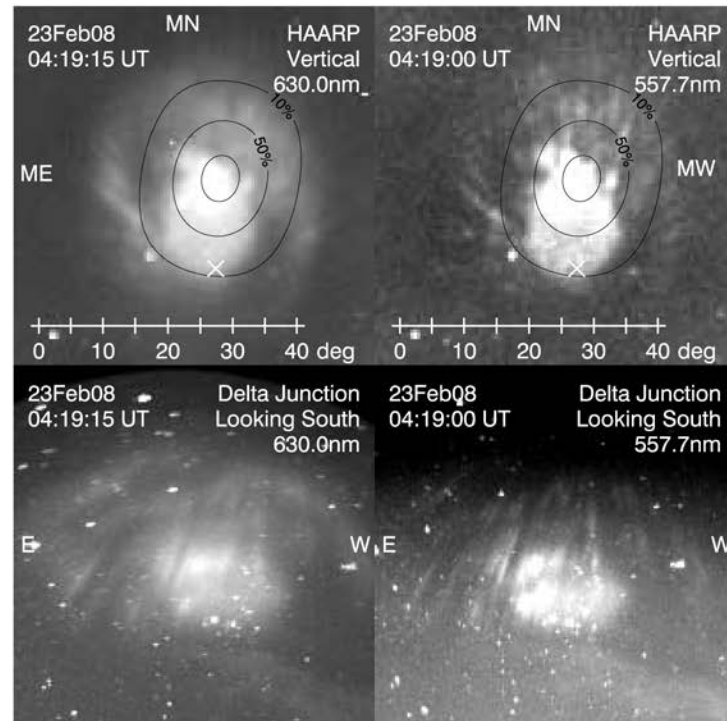


Observations and 3D simulations of effects from the constant heating of F-region

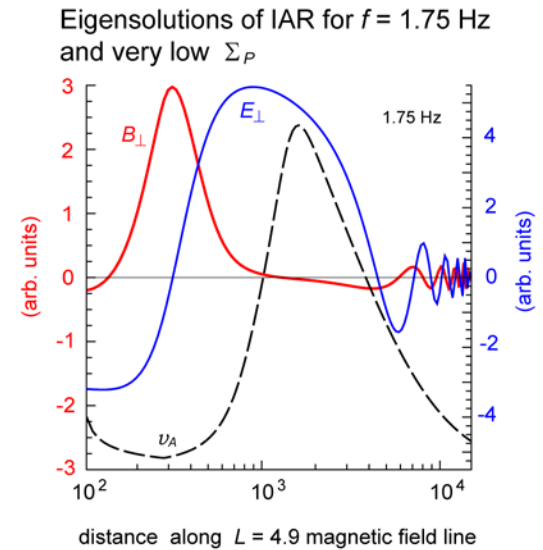
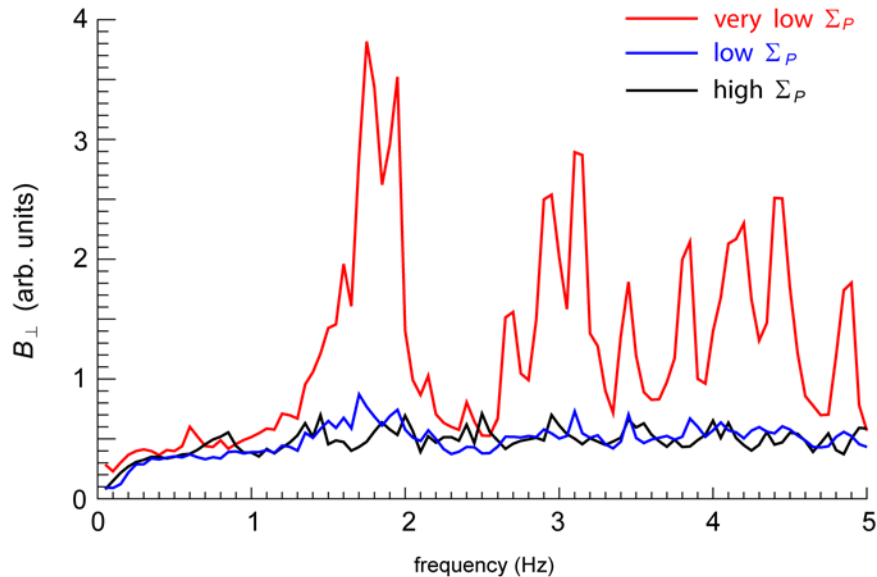
3D simulations of a constant heating of the same spot in the ionosphere (effect at 100 km altitude)



Observations of ring-like structures caused by the ionospheric heating [Pedersen et al, GRL, 2009]



Ionospheric
Alfvén
Resonator



Wave reflection from the ionosphere:

$$E_{\perp r} = R E_{\perp i}$$

$$R = \frac{\Sigma_A - \Sigma_P}{\Sigma_A + \Sigma_P}$$

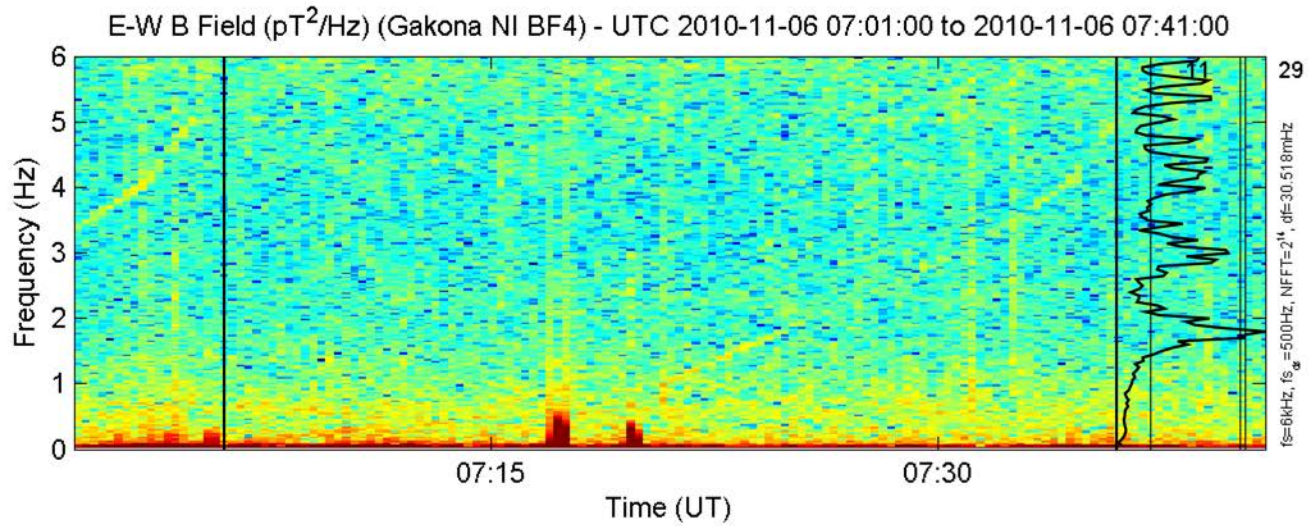
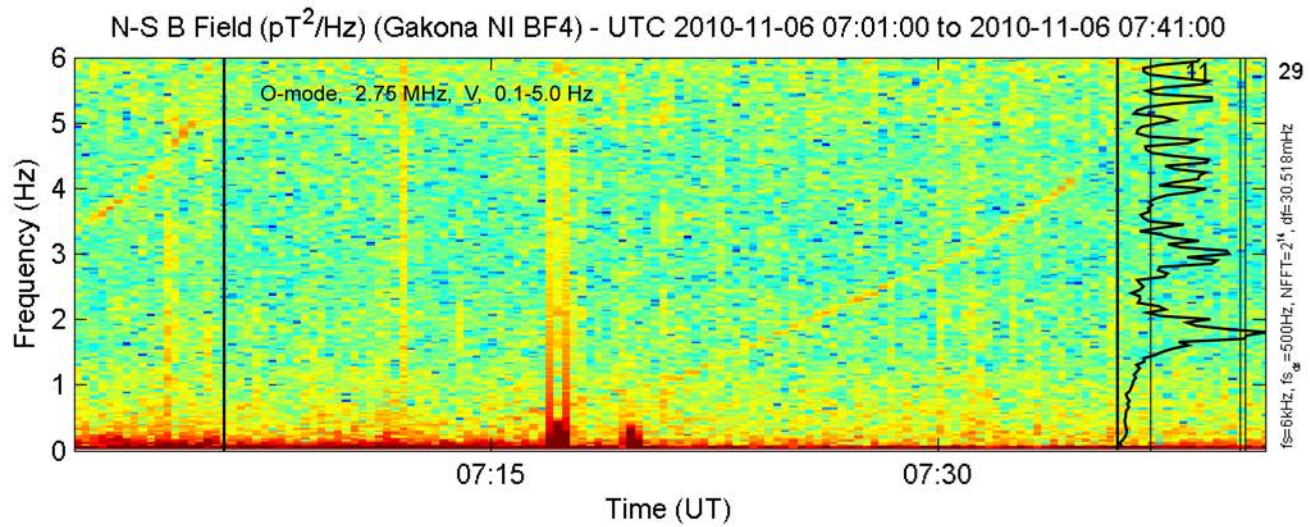
$$\Sigma_A = 1/\mu_0 v_A$$

high Σ_P : **$R = -0.310$**

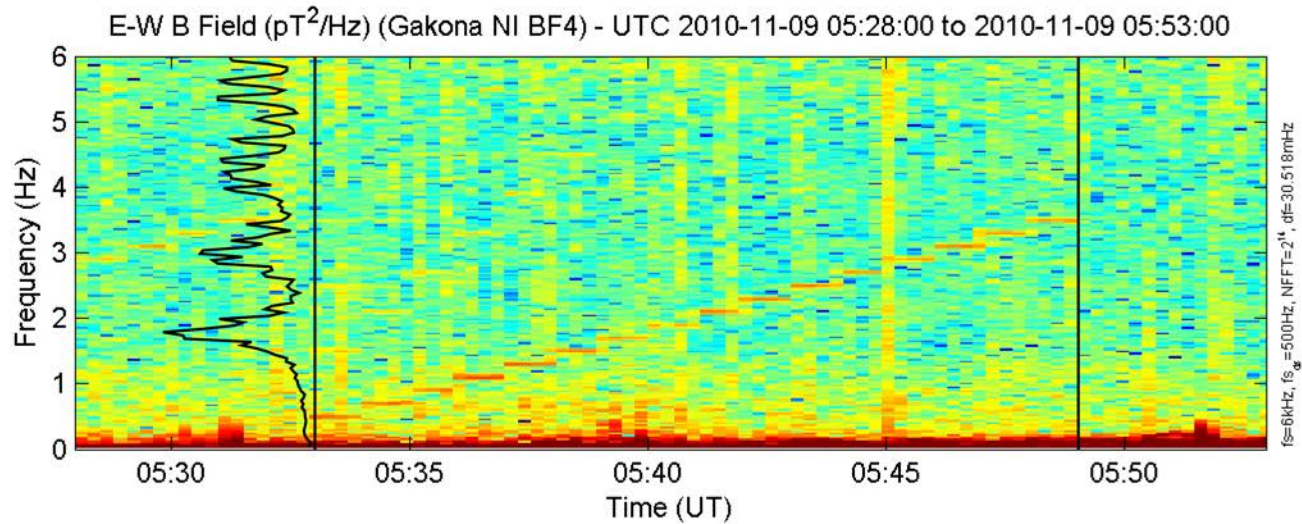
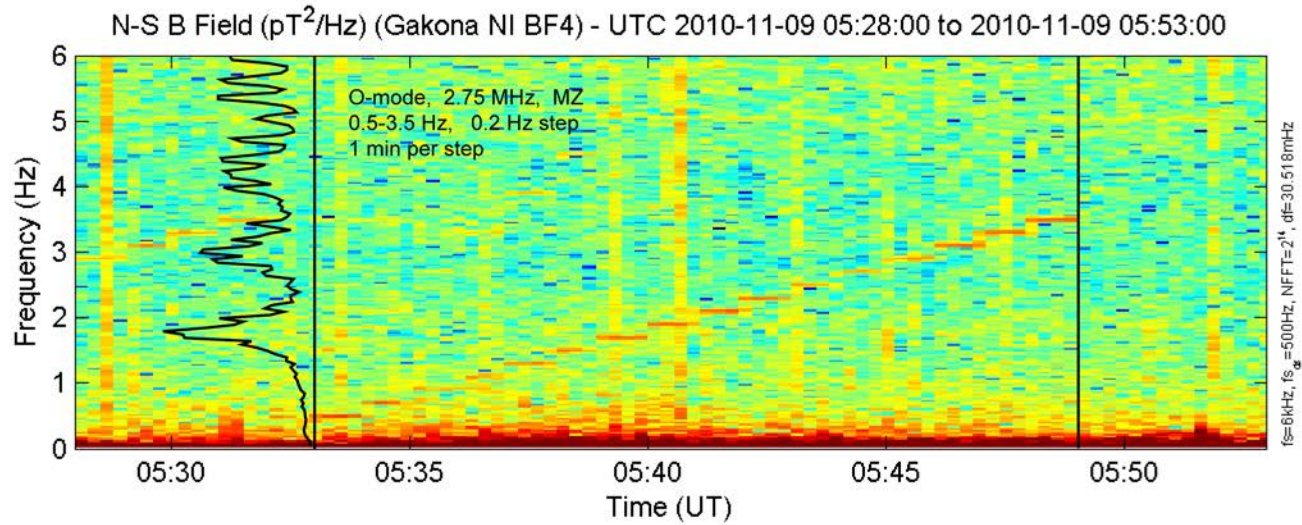
low Σ_P : **$R = 0.032$**

very low Σ_P : **$R = 0.510$**

Night (very low conductivity)

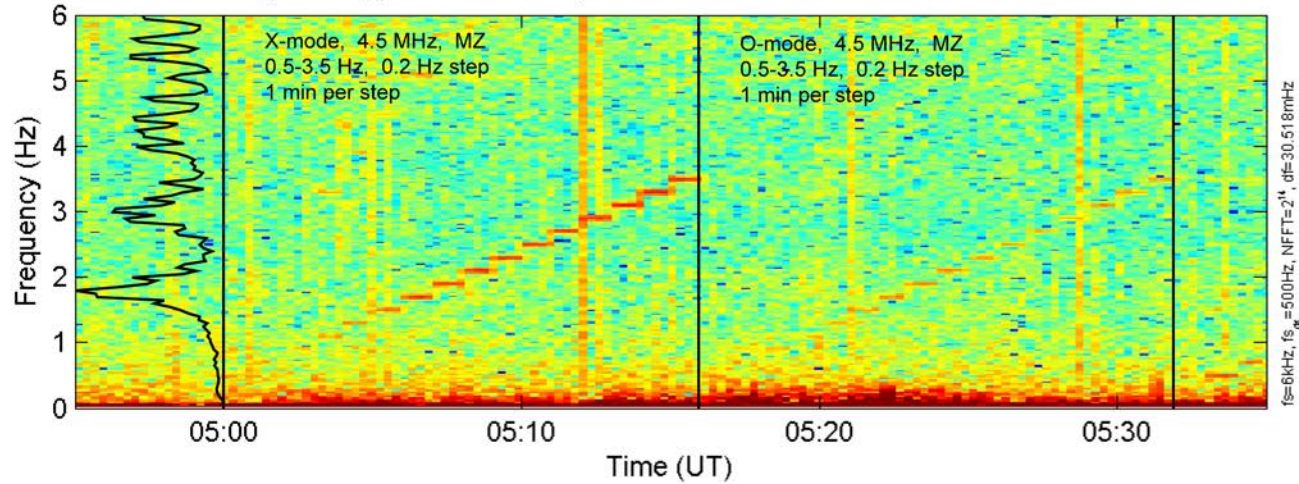


Night (very low conductivity)

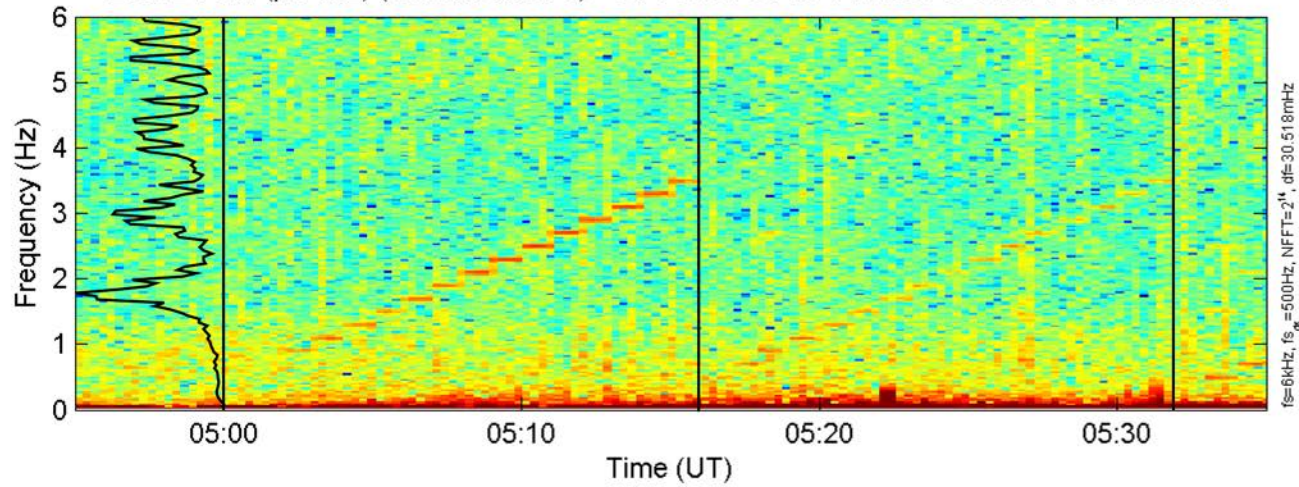


Night (very low conductivity)

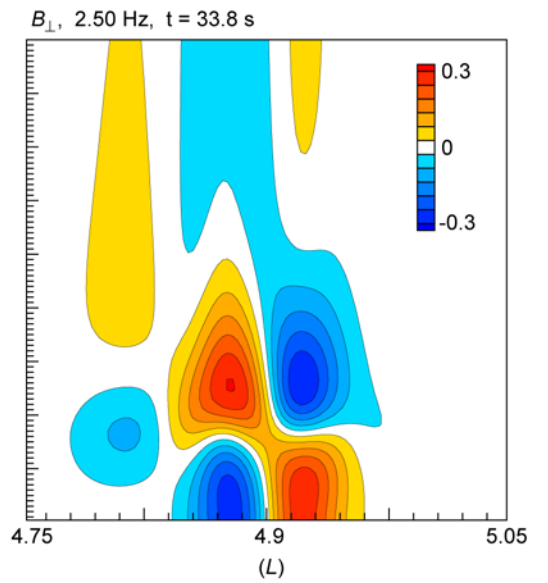
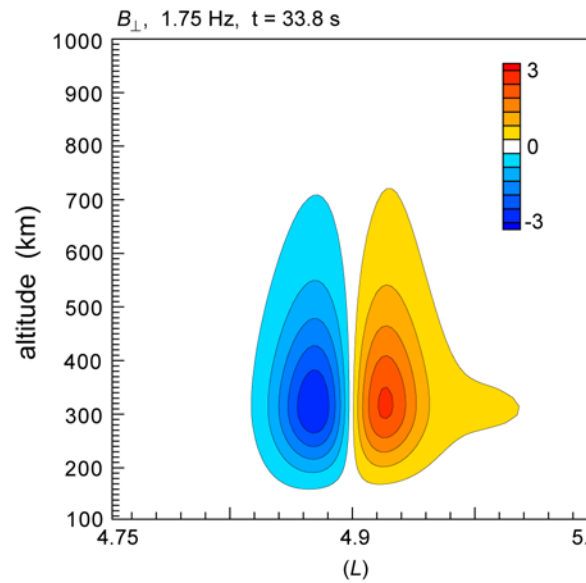
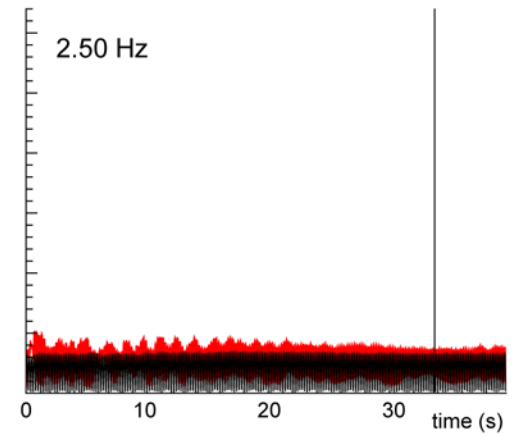
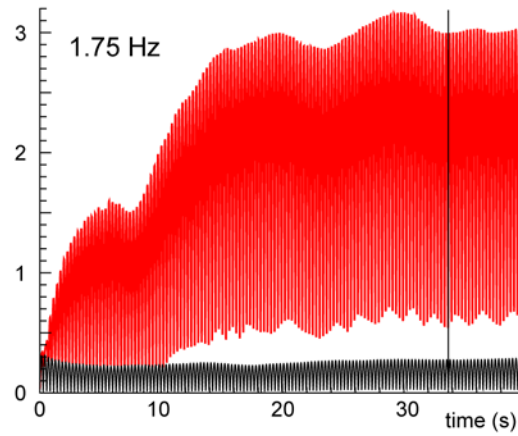
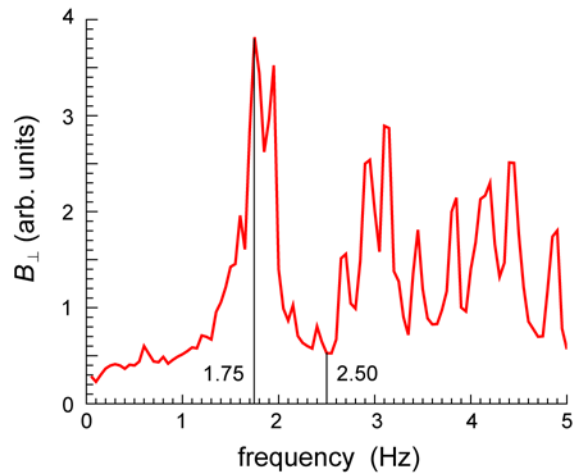
N-S B Field (pT^2/Hz) (Gakona NI BF4) - UTC 2010-11-09 04:55:00 to 2010-11-09 05:35:00



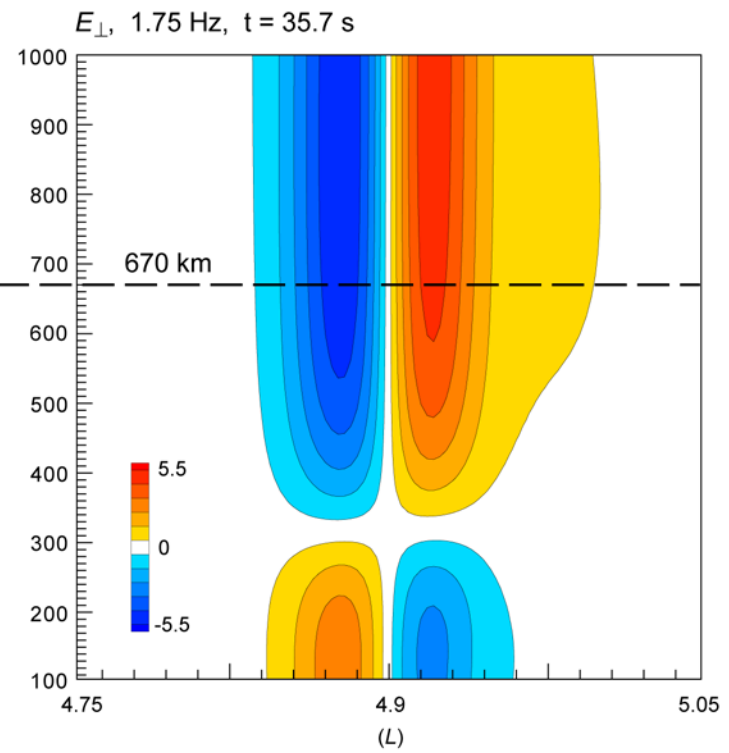
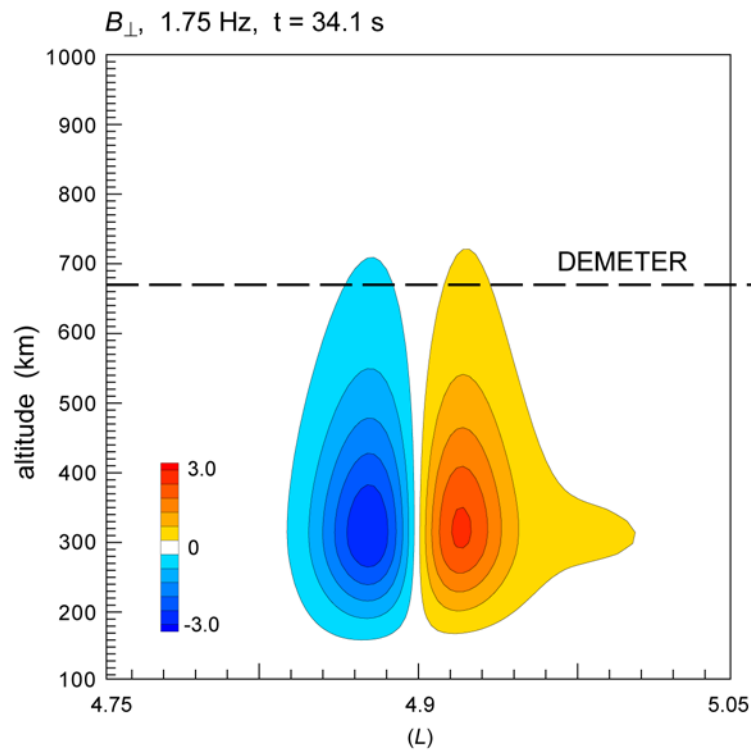
E-W B Field (pT^2/Hz) (Gakona NI BF4) - UTC 2010-11-09 04:55:00 to 2010-11-09 05:35:00



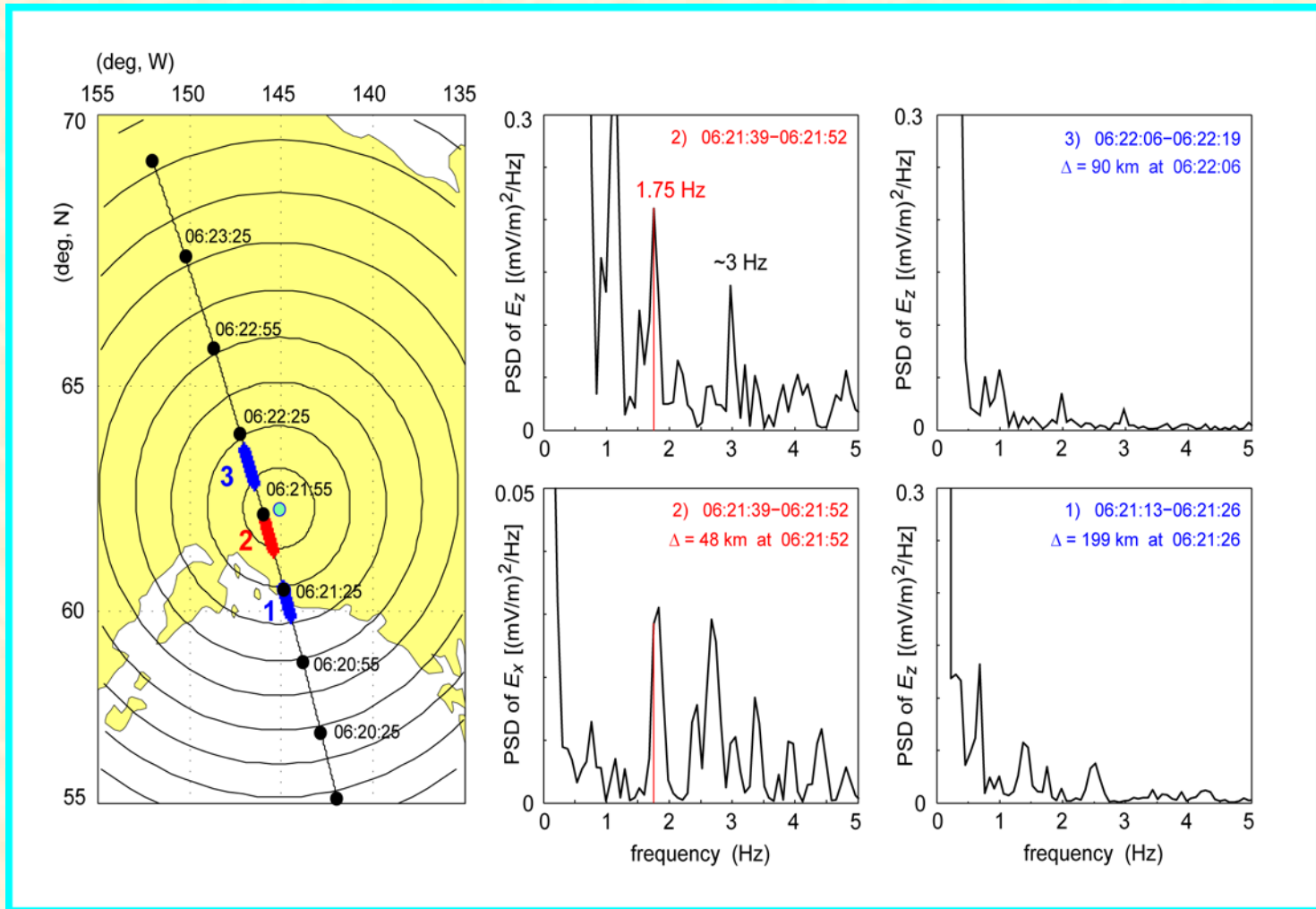
very low conductivity (night)



IAR, very low conductivity, 1.75 Hz



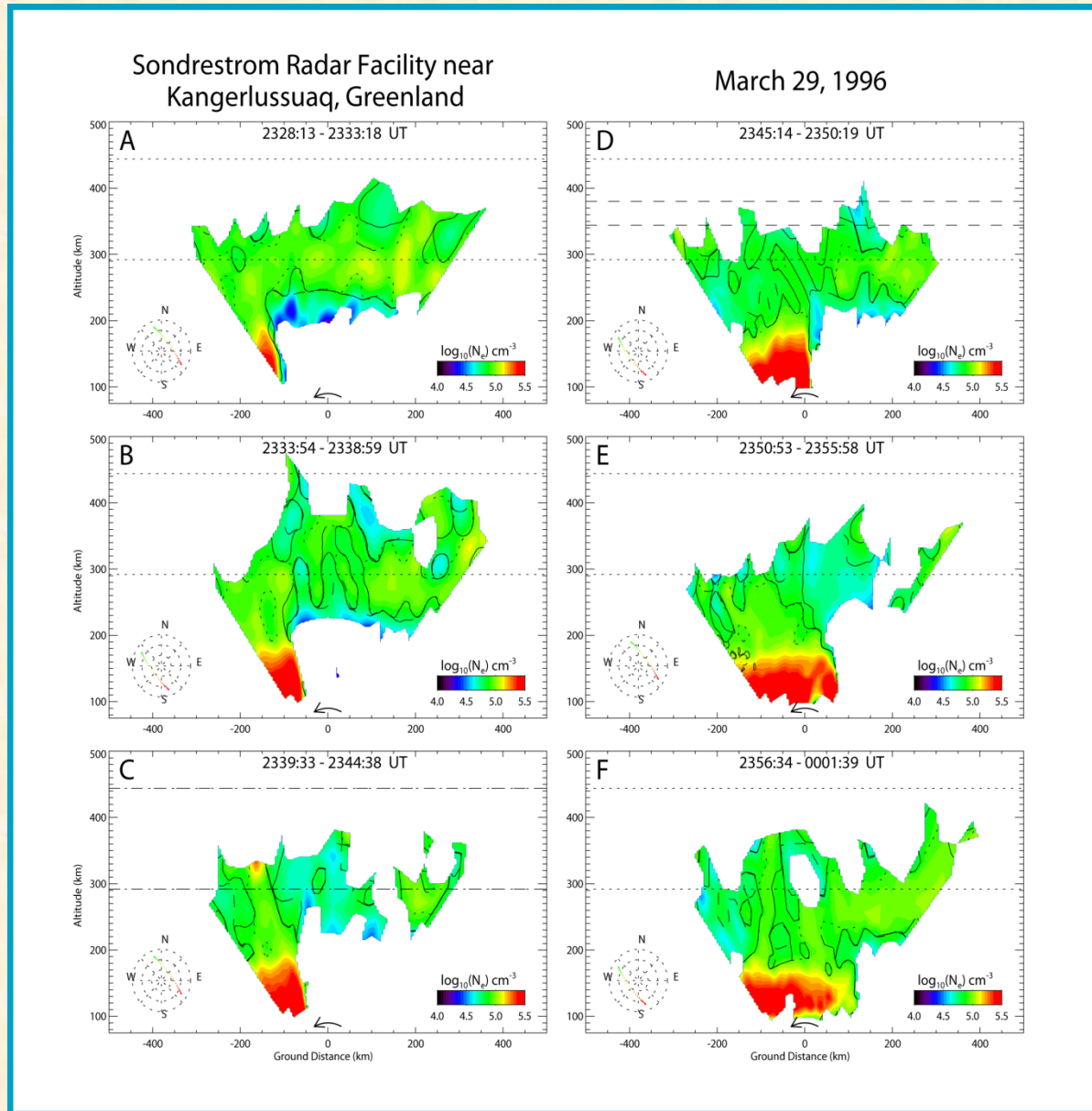
DEMETER - HAARP



Conclusions:

1. HAARP can efficiently generate large-amplitude ULF waves inside the global magnetospheric resonator and inside IAR.
2. The wave generation is most efficient when the ionospheric conductivity is very low (nighttime) and the heating is performed with X-mode waves in a frequency range from 2.8 to 4.5 MHz.
3. The structure of the waves inside IAR does NOT allow to make any conclusions about frequency of the resonator by measuring magnetic signals on the ground.
4. The resonant wave can be determined from measuring electric field on the ground or electric and magnetic field on satellites and/or sounding rockets.

Density cavities in a vicinity of auroral arcs



Density cavities in the downward current channel

